

Goal of the module: measure viscosity
using the Stokes-Einstein relation

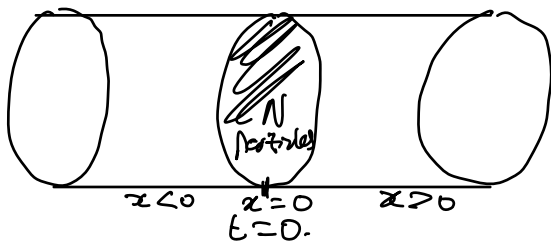
$$D = \frac{kT}{6\pi a \eta}$$

2 primary steps for development of the theory underpinning this module.

1. Einstein's analysis of Brownian motion
2. Stokes law for viscous drag on a sphere

Brownian Motion

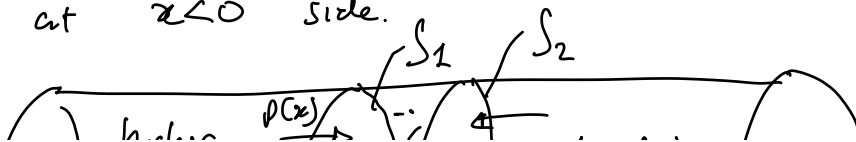
Thermal "kicks" due to molecular motion
kinetic motion of gas/liquid molecules causes
small particles to "feel" an unbalanced pressure
- This leads to diffuse motion.

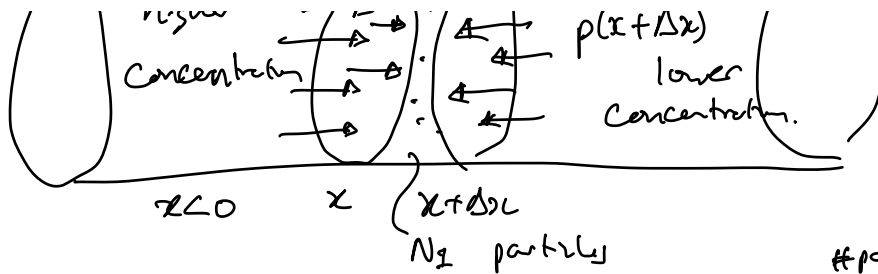


After some finite time, particles will move away from their initial positions..

Suppose that at some later time, more particles are located in the $x < 0$ side of the cylinder:

Then, an osmotic pressure, $p(x)$, will develop due to the larger concentration of particles at $x < 0$ side.





$$\frac{p(x) - p(x+\Delta x)}{\Delta x} = \frac{F(x) - F(x+\Delta x)}{A \cdot \Delta x} = \frac{\overset{\text{\# particles}}{N_1} \overset{\text{Force/particle}}{F_p}}{A \cdot \Delta x}$$

take limit as $\Delta x \rightarrow 0$:

$$\frac{\partial p}{\partial x} = - \underset{\substack{\uparrow \\ \text{Number density / concentration of} \\ \text{particles}}}{c} \cdot F_p$$

For every particle in motion at velocity v , a resistance or drag will supply a force in the opposite direction: F_f

$$F_f = m\beta v$$

Assuming the system is overdamped,

$$F_p = F_f \\ = m\beta v$$

Since particles in differential volume ΔV

satisfy the equation of state, $\#$ moles of Brownian particles

$$N_1 = n N_0 \quad \leftarrow \text{Avogadro's constant } \left(\frac{\# \text{ molecules}}{\text{mol}} \right)$$

$$\boxed{pV = n N_0 kT}$$

In the differential volume,

$$p \Delta V = N_1 kT \quad \text{so} \quad p = \underset{\substack{\uparrow \\ N_1 \\ \Delta V}}{c} kT$$

Differentiating: $\frac{\partial p}{\partial x} = kT \frac{\partial C}{\partial x} = -C F_p = -C m \beta v$

Using our expression, above, for F_p :

From Fick's law: # particles passing through area ΔA in time interval Δt is:

$$\Delta N = -D \frac{\partial C}{\partial x} \cdot \Delta A \cdot \Delta t \iff \frac{N_1}{A \cdot \Delta t} = -D \frac{\partial C}{\partial x}$$

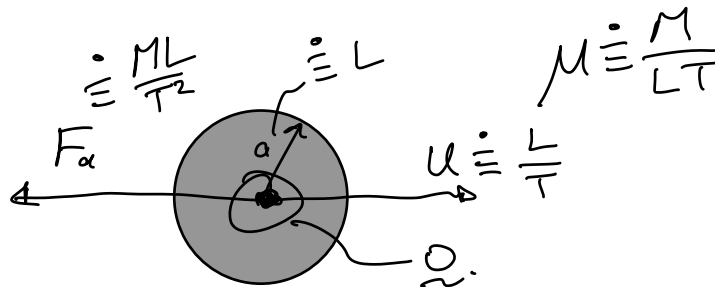
$$\implies C v = \frac{N_1}{A \cdot \Delta t} \cdot \frac{\Delta x}{\Delta t} = \frac{N_1}{A \Delta t} = -D \frac{\partial C}{\partial x}$$

Finally, we obtain:

$$\begin{aligned} kT \frac{\partial C}{\partial x} &= -m \beta C v \\ &= m \beta D \frac{\partial C}{\partial x} \implies \boxed{\frac{kT}{m\beta} = D} \end{aligned}$$

This relates diffn coeff to temp. & particle mobility. // Einstein.

Stokes drag on a sphere



from dimensional analysis: $F_d \propto a U \mu$.

$$\implies F_d = A a U \mu$$

constant is not yet determined.
From Stokes: 6π

Fluid motion is governed by Navier - Stokes:

$$\rho \frac{\partial u}{\partial t} = -\nabla p + \mu \nabla^2 u$$

For over-damped / low-Re flow, $\frac{D\underline{u}}{Dt} \rightarrow 0$:

This simplifies to:

$$\left. \begin{aligned} \nabla \left(\frac{p-p_0}{\mu} \right) = \nabla^2 \underline{u} = -\nabla \wedge \underline{\omega} \\ \text{incompressibility: } \nabla \cdot \underline{u} = 0. \end{aligned} \right\} \nabla^2 p = 0 \quad \& \quad \nabla^2 \underline{\omega} = 0.$$

If $c \rightarrow \infty$ the fluid is at rest $\&$ $p = p_0$,
and the particle moves at velocity \underline{U} :

Boundary conditions: $\underline{u} = \underline{U}$ on the sphere (no-slip condition)

$$\underline{u} \rightarrow 0, \quad p \rightarrow p_0 \quad \text{as } |\underline{x}| \rightarrow \infty$$

Note that the equations of motion $\&$ B.C.'s
are linear $\&$ homogeneous in \underline{u} , $\frac{p-p_0}{\mu}$, $\&$ \underline{U} ;
 $\Rightarrow \underline{u}$, $\frac{p-p_0}{\mu}$ are also linear $\&$ homogeneous in \underline{U} .

Solutions for \underline{u} $\&$ $\frac{p-p_0}{\mu}$ must be symmetrical
about the axis \parallel to \underline{U} .

$\Rightarrow \underline{u}$ must be in a plane through the axis
 $\parallel \underline{U}$.

Our differential operators are independent of
the choice of coordinates we use to describe them;
thus, our solution depends only on \underline{x} , and
not any other combination of its components.

\underline{u} $\&$ p can only depend on \underline{x} , \underline{U} , $\&$ a .

$$\Rightarrow \frac{p-p_0}{\mu} = \underline{U} \cdot \underline{x} f$$

where f is a function of $\frac{\underline{x} \cdot \underline{x}}{a^2} = \frac{r^2}{a^2}$ only.

Note that $p - p_0 = 0$ as $r \rightarrow \infty$ satisfies Laplace's eqn. \Rightarrow we can write $p - p_0$ as a series of spherical harmonics that scale inversely with r .

The only function that satisfies these conditions is the 2nd degree / dipole term:

$$\frac{p - p_0}{\mu} = C \frac{U \cdot x}{r^3}$$

The exact same reasoning applies for ω :

$$\omega = C \frac{U \wedge x}{r^3}$$

We know the azimuthal component of ω

$$\omega_\theta = \left[\frac{1}{r} \frac{\partial(rU_\theta)}{\partial r} - \frac{1}{r} \frac{\partial U_r}{\partial \theta} \right] \leftarrow \text{enforces incompressibility}$$

We can use the stream function, ψ , in spherical polar coordinates to express the velocity components:

(Note: I don't develop or derive these):

$$\left[U_r = \frac{1}{r^2 \sin \theta} \frac{\partial \psi}{\partial \theta} ; U_\theta = -\frac{1}{r \sin \theta} \frac{\partial \psi}{\partial r} \right]$$

$$\frac{\partial^2 \psi}{\partial r^2} + \frac{\sin \theta}{r^2} \frac{\partial}{\partial \theta} \left(\frac{1}{\sin \theta} \frac{\partial \psi}{\partial \theta} \right) = -\frac{C U \sin^2 \theta}{r} \quad (*)$$

We see that the particular solution for ψ is proportional to $\sin^2 \theta$.

The inner boundary condition requires that $\psi \propto \sin^2 \theta$

$$\text{at } r = a \text{ as well; } \Rightarrow \psi = U \sin^2 \theta f(r) \leftarrow$$

Using this expression for ψ , we can write

$$\underline{u} = \underline{U} \left(\frac{1}{r} \frac{df}{dr} \right) + \underline{x} \cdot \underline{U} \left(\frac{2f}{r^2} - \frac{1}{r} \frac{df}{dr} \right)$$

Canceling out $U \sin^2 \theta$ from \textcircled{A} above, we

find:

$$\boxed{\frac{d^2 f}{dr^2} - \frac{2f}{r^2} = -\frac{C}{r}}$$

$$\Rightarrow \boxed{f(r) = \frac{Cr}{2} + \frac{L}{r} + Mr^2}$$

The B.C. on f as $r \rightarrow \infty$:

$$\boxed{\frac{f}{r^2} \rightarrow 0}$$

no-slip $\Rightarrow u_r = U \cos \theta$ for

$$r=a \Leftrightarrow$$

$$\boxed{f(a) = \frac{a^2}{2}}$$

$$\Rightarrow M=0 \Leftrightarrow L = \frac{a^3}{2} - \frac{Ca^2}{2}$$

$$u_\theta \Big|_{r=a} = \left[-\frac{1}{r \sin \theta} \frac{\partial \psi}{\partial r} \right]_{r=a} = -U \sin \theta \Big|_{r=a} \Rightarrow C = \frac{3a}{2}$$

$$\text{Thus, } \boxed{\psi = U r^2 \sin^2 \theta \left(\frac{3}{4} \frac{a}{r} - \frac{1}{4} \frac{a^3}{r^3} \right)}$$

The force acting to resist the sphere's motion is equal to the integral of the stress over the sphere's surface:

$$n_j (\sigma_{ij})_{r=a} = n_j \left\{ -p \delta_{ij} + \mu \left(\frac{\partial u_i}{\partial x_j} + \frac{\partial u_j}{\partial x_i} \right) \right\}_{r=a}$$

With some algebra, for our velocity field, we find

$$n_j (\sigma_{ij})_{r=a} = \left\{ -p n_i + \mu n_i \underline{U} \cdot \underline{n} \left(-\frac{f''}{r} + \frac{6f'}{r^2} - \frac{10f}{r^3} \right) + \mu n_i \left(\frac{f''}{r} - \frac{2f'}{r^2} + \frac{2f}{r^3} \right) \right\}_{r=a}$$

$$\text{note: } f' = \frac{df}{dr}$$

Substituting in for p & f

$$n_i (\sigma_{ij})_{r=a} = n_i \left\{ -p_0 + \frac{3\mu U \cdot n_j}{a} \left(\frac{2C}{a} - 3 \right) \right\} + \frac{3\mu U_i}{a} \left(1 - \frac{C}{a} \right)$$

$$= -p_0 n_i - \frac{3\mu U_i}{2a}$$

The first term is the same pressure at $r=\infty$.

\Rightarrow no net effect on the sphere

We can see that the stress is \parallel to particle motion

$$F_a = \int \underline{\sigma} \cdot \underline{n} dA = \frac{3\mu U}{2a} \cdot \underbrace{\int dA}_{4\pi a^2} = \boxed{6\pi a \mu U}$$

= F_p from Einstein's development

$$\Rightarrow m \beta = 6\pi a \mu$$

$$\Rightarrow \left| D = \frac{kT}{6\pi a \mu} \right|$$

↑
measured.

Stokes-Einstein